

All force directions toward center, or, " $-\hat{r}$ "

$$(a) \vec{F} = -\frac{G(M_1 + M_2)m}{a^2} \hat{r}$$

since shells uniform and all mass at radii $< a$, as if all mass concentrated at origin.

$$(b) \vec{F} = -\frac{GM_1 m}{b^2} \hat{r}$$

$$(c) \vec{F} = 0 \text{ no mass inside}$$



radius of hole is $-R/2$

\therefore volume is $\frac{1}{8}$ of big sphere's volume. $\dots (\frac{1}{2})^3$

use concept of superposition...

"LIKE" full sphere of radius

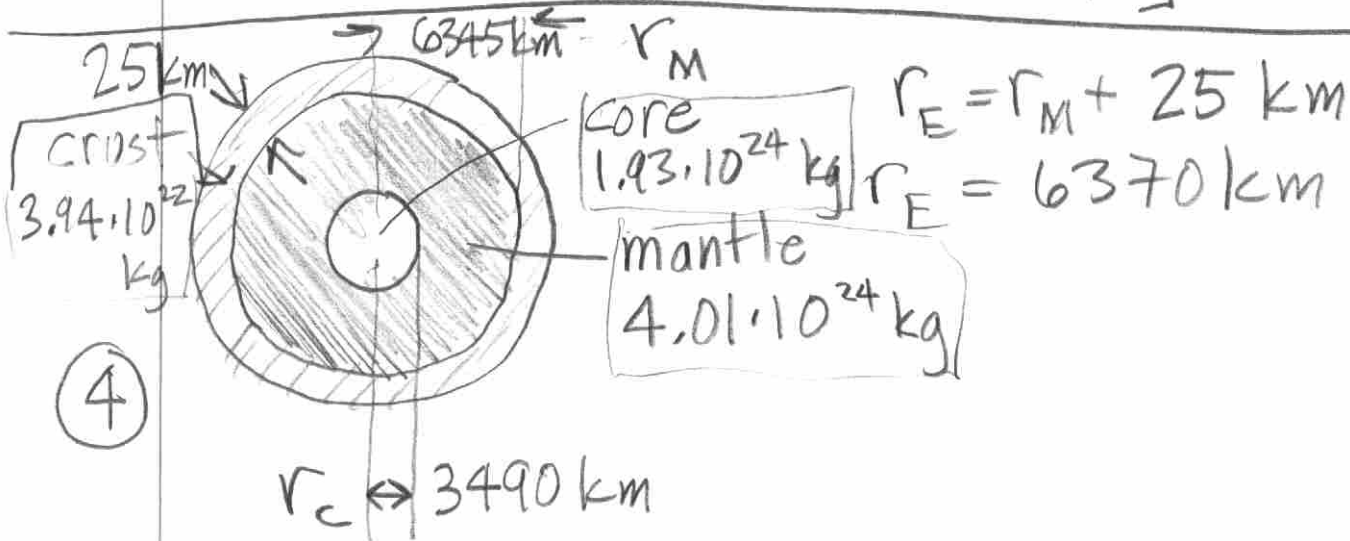
$R +$ mass sphere of $\frac{1}{8}$ mass at $R/2$

④
② after

$$\vec{F} = \left(-\frac{GMm}{d^2} + \frac{G\left(\frac{M}{8}\right)m}{(d - R/2)^2} \right) \hat{r}$$

↑ unit away from center

$$\vec{F} = -\frac{GMm}{d^2} \left[1 - \frac{1}{8(1 - R/2d)^2} \right] \hat{r}$$



④

(a) $M_{\text{tot}} = M_{\text{crust}} + M_{\text{mantle}} + M_{\text{core}}$

$$= (0.0394 + 4.01 + 1.93) \cdot 10^{24} \text{ kg}$$

$$= 5.98 \cdot 10^{24} \text{ kg}$$

$$g = \frac{GM_{\text{tot}}}{R_E^2} = \frac{(6.673 \cdot 10^{-11})(5.98 \cdot 10^{24})}{(6370 \cdot 10^3)^2}$$

$$g = 9.83 \text{ m/s}^2$$

(b) $M_{\text{inside}} = M_{\text{mantle}} + M_{\text{core}}$

$$= (4.01 + 1.93) \cdot 10^{24} \text{ kg}$$

$$= 5.94 \cdot 10^{24} \text{ kg}$$

$$R = r_m = 6345 \text{ km}$$

$$g = \frac{GM_{\text{inside}}}{r_m^2} = \frac{(6.673 \cdot 10^{-11})(5.94 \cdot 10^{24})}{(6345 \cdot 10^3)^2}$$

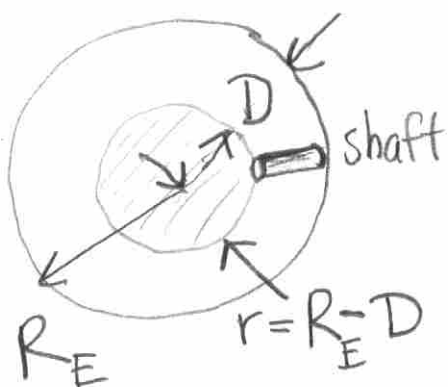
$$g = 9.85 \text{ m/s}^2$$

$$(c) \quad g = g_s \left(1 - \frac{D}{R_E}\right)$$

$$= 9.83 \left(1 - \frac{25}{6370}\right)$$

$$g = 9.79 \text{ m/s}^2$$

$$(2) \quad \rho_E = \frac{M_E}{V_E} = \frac{M_E}{\frac{4\pi}{3} R_E^3}$$



Mass inside D is:

$$M_{\text{inside}} = \frac{4\pi}{3} \rho_E (R_E - D)^3$$

$$g = \frac{GM_{\text{inside}}}{r^2}$$

$$= \frac{4\pi}{3} G \rho_E \frac{(R_E - D)^3}{(R_E - D)^2}$$

$$= \frac{4\pi}{3} G \frac{M_E}{\frac{4\pi}{3} R_E^3} (R_E - D) = \frac{GM_E}{R_E^2} \left(1 - \frac{D}{R_E}\right)$$

$$g = g_s \left(1 - \frac{D}{R_E}\right)$$

5. a

Boundary	\tilde{r} (boundary, km)	$\rho < (10^3 \frac{\text{kg}}{\text{m}^3})$	$\rho > (10^3 \frac{\text{kg}}{\text{m}^3})$
Core / Mantle	3490	10.8	4.50
Mantle / Crust	6345	5.55	3.10
Crust / Air	6375	5.52	≈ 0

This is

$$\frac{\sum M_i (\text{inside})}{\frac{4\pi}{3} r_{\text{boundary}}^3}$$

This is

$$\frac{M_i (\text{just outside})}{\frac{4\pi}{3} (r_{\text{next boundary}}^3 - r_{\text{boundary}}^3)}$$

b) The mass inside radius r has two contributions, consisting of

(i) mass inside $\tilde{r} = \frac{4\pi}{3} \rho < \tilde{r}^3$

(ii) mass between \tilde{r} and r :

$$= \frac{4\pi}{3} \rho > (r^3 - \tilde{r}^3)$$

$$g = \frac{G}{r^2} M_{\text{inside}} = \frac{G}{r^2} \left(\frac{4\pi}{3} \rho < \tilde{r}^3 + \frac{4\pi}{3} \rho > (r^3 - \tilde{r}^3) \right)$$

$$g = G \left[\frac{4\pi}{3} \frac{(\rho_{<} - \rho_{>}) \tilde{r}^3}{r^2} + \frac{4\pi}{3} \rho_{>} r \right]$$

$$\text{so } A = \frac{4\pi}{3} (\rho_{<} - \rho_{>}) \tilde{r}^3 \quad B = \frac{4\pi}{3} \rho_{>} r$$

but we know from problem 4 that when $\rho_{<} = \rho_{>} = \rho$ $g = \frac{4\pi}{3} G \rho r$, so, when ρ uniform, A must be zero, and it is.

$$(c) \quad g = \frac{G M_{\text{inside}}}{r^2}$$

$$M_{\text{inside}} = 1.93 \cdot 10^{24} \text{ kg (core)}$$

$$r = 3490 \text{ km (core)}$$

$$g_c = \frac{(6.673 \cdot 10^{-11}) (1.93 \cdot 10^{24})}{(3490 \cdot 10^3)^2}$$

$$g_c = 10.6 \text{ m/s}^2$$

at
core/
mantle
boundary

from $r=0$ to $r=r_c$ (core mantle boundary)

$$M_{\text{inside}} = \rho_{\text{core}} \cdot \frac{4\pi}{3} r^3 = \frac{M_{\text{core}}}{\frac{4\pi}{3} r_c^3} \cdot \frac{4\pi}{3} r^3 = M_{\text{core}} \left(\frac{r}{r_c}\right)^3$$

$$g(r) = \frac{G M_{\text{inside}}}{r^2} = \frac{G M_{\text{core}}}{r_c^2} \left(\frac{r}{r_c}\right)^3 = \frac{G M_{\text{core}}}{r_c^2} \left(\frac{r}{r_c}\right)$$

$$g(r) = g_c \left(\frac{r}{r_c}\right) = (10.6 \text{ m/s}^2) \left(\frac{r}{3490 \text{ km}}\right)$$

$$(d) \quad g = G \left(\frac{A}{r^2} + Br \right)$$

$$\frac{dg}{dr} = G \left(-\frac{2A}{r^3} + B \right) = 0$$

$$r_{\min} = \left(\frac{2A}{B} \right)^{1/3} = \left(2 \left[\frac{\rho_{<}}{\rho_{>}} - 1 \right] \right)^{1/3} \tilde{r}$$

$$g_{\min} = G \left(\frac{A}{\left(\frac{2A}{B} \right)^{2/3}} + B \left(\frac{2A}{B} \right)^{1/3} \right)$$

$$= G \left(\frac{1}{2^{2/3}} (B^2 A)^{1/3} + 2^{1/3} (B^2 A)^{1/3} \right)$$

$$g_{\min} = G \left(\frac{1}{2^{2/3}} + 2^{1/3} \right) (AB^2)^{1/3}$$

$$= \frac{4\pi}{3} G \left(\frac{1}{2^{2/3}} + 2^{1/3} \right) \left((\rho_{<} - \rho_{>}) \rho_{>}^2 \tilde{r}^3 \right)^{1/3}$$

$$g_{\min} = \frac{4\pi}{3} G \rho_{>} \tilde{r} \left(\frac{1}{2^{2/3}} + 2^{1/3} \right) \left(\frac{\rho_{<}}{\rho_{>}} - 1 \right)^{1/3}$$

$$\frac{d^2g}{dr^2} = G \cdot \frac{6A}{r_{\min}^4} > 0, \quad \text{above is a minimum;}$$

$$\text{need } \frac{\rho_{<}}{\rho_{>}} > 1,$$

higher density inside.

for $r_{\min} > \tilde{r}$, need

$$\left(2 \left[\frac{\rho_{<}}{\rho_{>}} - 1 \right] \right)^{1/3} > 1 \Rightarrow \left(\frac{\rho_{<}}{\rho_{>}} > \frac{3}{2} \right)$$

numerically,

$$r_{\min} = \left(2 \cdot \left(\frac{10.8}{4.50} - 1 \right) \right)^{1/3} \cdot 3490 \text{ km}$$

$$r_{\min} = 4932 \text{ km}$$

$$g_{\min} = \frac{4\pi}{3} \cdot (6.673 \cdot 10^{-11}) (4.50 \cdot 10^3) (3490 \cdot 10^3) \\ \times \left(\frac{1}{2^{2/3}} + 2^{1/3} \right) \left(\frac{10.8}{4.50} - 1 \right)^{1/3}$$

$$g_{\min} = 9.30 \text{ m/s}^2$$

(e) core: $g(r) = \left(10.6 \frac{\text{m}}{\text{s}^2} \right) \left(r / 3490 \text{ km} \right)$

mantle: $g(r) = \frac{4\pi}{3} G(\rho_c - \rho_m) \tilde{r} \cdot \left(\frac{\tilde{r}}{r} \right)^2 + \frac{4\pi}{3} G \rho_m \tilde{r} \left(\frac{r}{\tilde{r}} \right)$

$$= \frac{4\pi}{3} \cdot (6.673 \cdot 10^{-11}) (10.8 - 4.5) \cdot 10^3 \cdot (3490 \cdot 10^3)$$

$$= 6.19 \text{ m/s}^2$$

$$\downarrow = \frac{4\pi}{3} (6.673 \cdot 10^{-11}) (4.5 \cdot 10^3) \cdot (3490 \cdot 10^3)$$

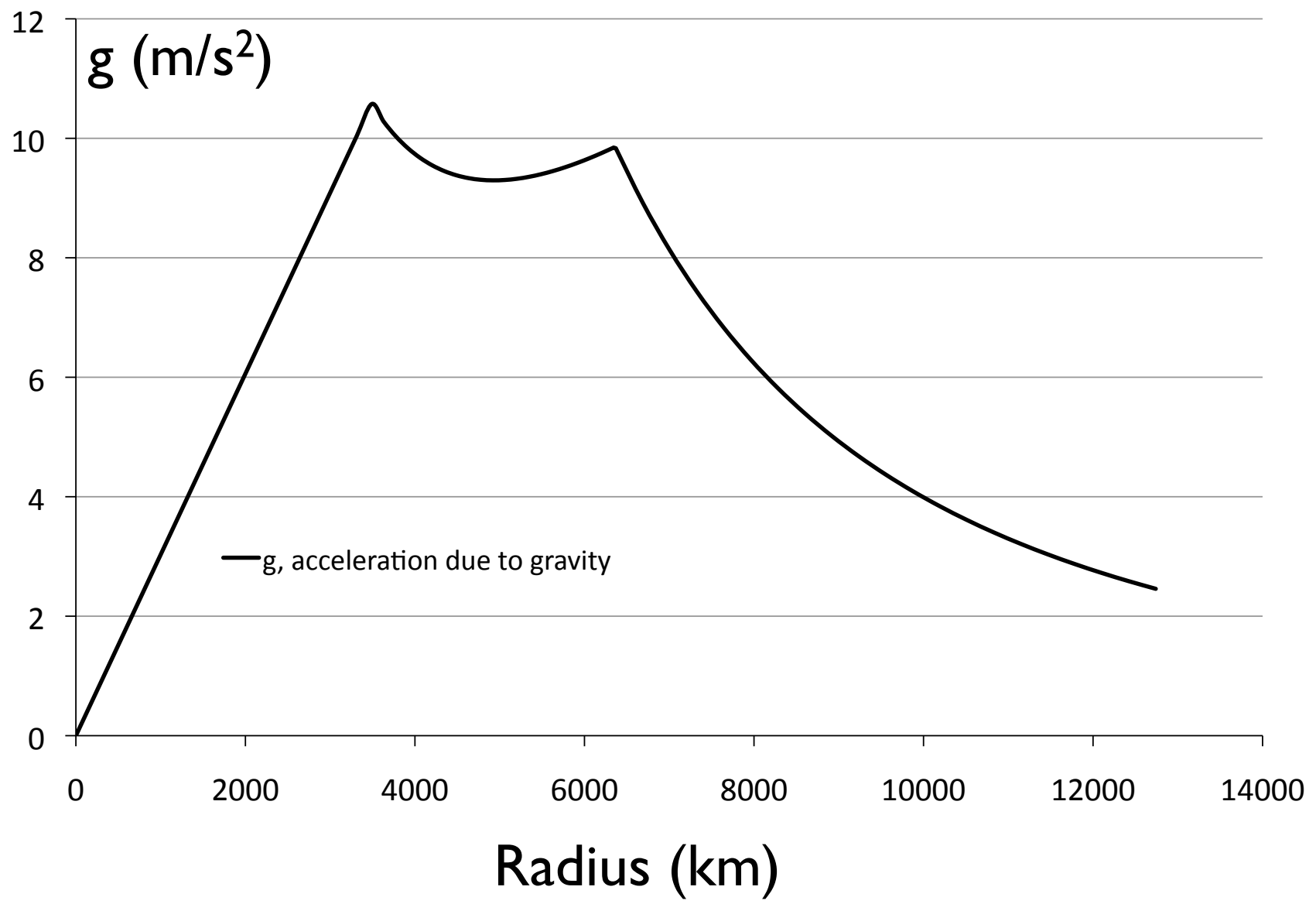
$$= 4.39 \text{ m/s}^2$$

$$g(r) = \left(6.19 \frac{\text{m}}{\text{s}^2} \right) \left(\frac{3490 \text{ km}}{r} \right)^2 + \left(4.39 \frac{\text{m}}{\text{s}^2} \right) \left(\frac{r}{3490 \text{ km}} \right)$$

crust: $\frac{4\pi}{3} G(\rho_c - \rho_m) \tilde{r} = \frac{4\pi}{3} (6.673 \cdot 10^{-11}) (5.55 - 3.10) \cdot 10^3 (6345 \cdot 10^3)$

$$= 1.24 \text{ m/s}^2$$

Effect of core, mantle, crust density on acceleration of gravity in and above the Earth



$$\frac{4\pi}{3} G \rho \vec{r} = \frac{4\pi}{3} (6.673 \cdot 10^{-11}) (3 \cdot 10^3) \cdot (6345 \cdot 10^3)$$

$$= 5.50 \text{ m/s}^2$$

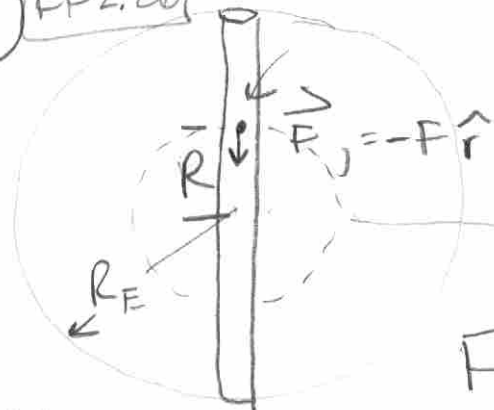
$$g(r) = \left(4.34 \frac{\text{m}}{\text{s}^2}\right) \left(\frac{6345 \text{ km}}{r}\right)^2 + \left(5.50 \frac{\text{m}}{\text{s}^2}\right) \left(\frac{r}{6345 \text{ km}}\right)$$

air :

$$g = \frac{GM_E}{r^2} = \underbrace{\frac{GM_E}{R_E^2}}_{g_s} \cdot \left(\frac{R_E}{r}\right)^2$$

$$g(r) = \left(9.83 \frac{\text{m}}{\text{s}^2}\right) \left(\frac{6370 \text{ km}}{r}\right)^2$$

6) KX 2.26



$$\rho_E = \frac{M_E}{\frac{4\pi}{3} R_E^3}$$

$$M_{\text{inside}} = \frac{4\pi}{3} \rho_E R^3$$

$$F = G \frac{M_{\text{inside}} m}{R^2}$$

$$= G \frac{\frac{4\pi}{3} \rho_E R^3 \cdot m}{R^2}$$

$$= \left(G \frac{4\pi}{3} \rho_E R_E\right) \left(\frac{R}{R_E}\right) \cdot m$$

$$= \underbrace{G \frac{M_E}{R_E^2}}_g \cdot \left(\frac{R}{R_E}\right) \cdot m$$

g

$$F = \left(\frac{R}{R_E}\right) mg = \underbrace{\left(\frac{mg}{R_E}\right)}_{\text{"k"}} R$$

direction is $-$,

$$\vec{F} = -kR \hat{r}$$

$$\omega = \sqrt{\frac{k}{m}} = \sqrt{\frac{mg}{mR_E}} = \sqrt{\frac{g}{R_E}}$$

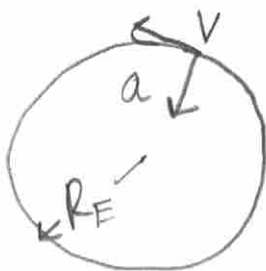
$$T = \frac{2\pi}{\omega} = 2\pi \sqrt{\frac{R_E}{g}}$$

$$= 2\pi \cdot \sqrt{\frac{6370 \cdot 10^3}{9.83}}$$

$$T = 5060 \text{ s} = 84.3 \text{ min}$$

$$= 1 \text{ hr } 24 \text{ minutes}$$

Low Earth Orbit



$$F_{\text{net}} = ma$$

$$g = \frac{GM_E m}{R_E^2} = m \cdot \frac{v^2}{R_E}$$

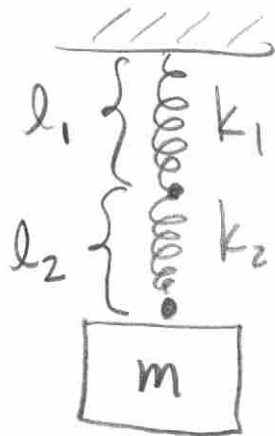
$$v \cdot T = 2\pi R_E$$

$$T = \frac{2\pi R_E}{v}$$

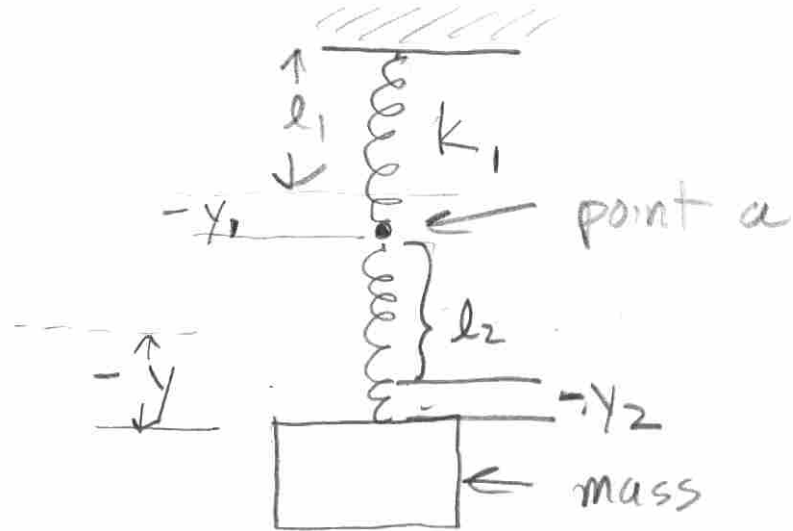
$$\text{so } v = \sqrt{gR_E}$$

$$T = \frac{2\pi R_E}{v} = \frac{2\pi R_E}{\sqrt{gR_E}} = 2\pi \sqrt{\frac{R_E}{g}} \quad \underline{\underline{\text{same!}}}$$

7. KK 2.31

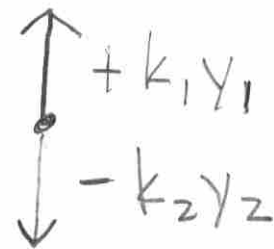


(a)



$$-y = -y_1 - y_2$$

point a - no mass,



$$k_1 y_1 - k_2 y_2 = 0$$

$$y_1 = \frac{k_2}{k_1} y_2$$

so $y = y_1 + y_2 = \left(\frac{k_2}{k_1} + 1\right) y_2$

but on mass $y_2 = \frac{y}{\left(\frac{k_2}{k_1} + 1\right)}$

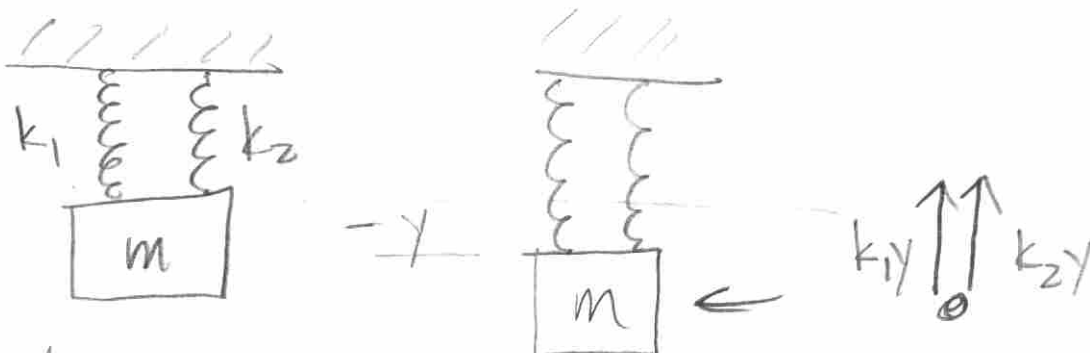


$$F_{\text{springs}} = k_2 y_2 = \frac{k_2}{\frac{k_2}{k_1} + 1} y$$

$$F_{\text{springs}} = \frac{1}{\frac{1}{k_1} + \frac{1}{k_2}} y = k_a y$$

or $\frac{1}{k_a} = \frac{1}{k_1} + \frac{1}{k_2}$, $k_a = \frac{k_1 k_2}{k_1 + k_2}$

$$\omega_a = \sqrt{\frac{k_a}{m}} = \sqrt{\frac{1}{m} \frac{k_1 k_2}{k_1 + k_2}}$$



(b)

$$F = \underset{\text{springs}}{(k_1 + k_2)y} = k_b y$$

$$k_b = k_1 + k_2$$

$$\omega_b = \sqrt{\frac{k_b}{m}} = \sqrt{\frac{k_1 + k_2}{m}}$$

Hint: $k_1 = k_2 = k$

$$\omega_a = \sqrt{\frac{1}{m} \frac{k^2}{k+k}} = \sqrt{\frac{k}{2m}} \quad \checkmark$$

$$\omega_b = \sqrt{\frac{1}{m} \cdot 2k} = \sqrt{\frac{2k}{m}} \quad \checkmark$$