

# Quantum Simple Harmonic Oscillator

$$\frac{p^2}{2m} + \frac{1}{2}(k = m\omega^2)x^2 |E\rangle = E |E\rangle$$

⇓ rep. in  $|x\rangle$

$$\left( \frac{-\hbar^2}{2m} \frac{d^2}{dx^2} + \frac{1}{2} m\omega^2 x^2 \right) \psi(x) = E \psi(x)$$

⇓

$$y \equiv bx$$

$$b \equiv \sqrt{\frac{\hbar}{m\omega}}$$

$$\epsilon \equiv \frac{E}{\hbar\omega}$$

$$\frac{d^2\psi}{dy^2} + (2\epsilon - y^2)\psi = 0$$

Split this into two limiting cases:

(I)  $y^2 \gg 2\epsilon$

$$\frac{d^2\psi}{dy^2} - y^2\psi = 0$$

usual trick:  $\psi = A e^{f(y)}$

$$\frac{d\psi}{dy} = A f'(y) e^{f(y)}$$

$$\frac{d^2\psi}{dy^2} = A (f''(y) + (f'(y))^2) e^{f(y)}$$

so want

$$A (f''(y) + [f'(y)]^2 - y^2) e^{f(y)} = 0$$

a pretty obvious ansatz:

$$[f'(y)]^2 = y^2$$

$$f'(y) = \pm y \Rightarrow \underline{f(y) = \pm \frac{1}{2}y^2} + \text{(constant absorbed into A)}$$

then

$$\star A (\pm 1 + \underbrace{y^2 - y^2}_{\text{highest order in } y, \text{ cancelling}}) e^{\pm \frac{1}{2}y^2} = 0$$

relatively negligible as  $y \rightarrow \infty$   
(as justified in ignoring as we were in ignoring  $2\epsilon$ !)

Turns out, at this level of accuracy, a more general  $\psi$  works too:

$$\psi(y) = Ay^m e^{\pm y^2/2}$$

$$\frac{d\psi}{dy} = Amy^{m-1} e^{\pm y^2/2} \pm Ay^{m+1} e^{\pm y^2/2}$$

$$\frac{d^2\psi}{dy^2} = \underbrace{(Am(m-1)y^{m-2} \pm 2Amy^m)}_{\text{negligible}} + Ay^{m+2} e^{\pm y^2/2}$$

$$y^2\psi = - ( \text{negligible terms} + Ay^{m+2} e^{\pm y^2/2} )$$

these terms, as  $y \rightarrow \infty$ , are negligible relative to  $y^{m+2}$

$= 0$  (to same approx as ignoring  $2\epsilon$  gave).

(II)  $y \rightarrow 0$

$$\frac{d^2\psi}{dy^2} + 2\varepsilon\psi = 0 \quad \psi = \alpha e^{f(y)}$$
$$\alpha (f''(y) + [f'(y)]^2 + 2\varepsilon) e^{f(y)} = 0$$

$\rightarrow 0$

$$f'(y) = \pm \sqrt{2\varepsilon} i$$
$$f(y) = \pm \sqrt{2\varepsilon} i y \quad (+ \text{constant absorbed in } \alpha)$$

$$\psi(y) = \alpha e^{+\sqrt{2\varepsilon} i y} + \beta e^{-\sqrt{2\varepsilon} i y}$$

as  $y \rightarrow 0$   $\psi(y) \rightarrow \underbrace{(\alpha + \beta)}_A + \underbrace{(\alpha - \beta)\sqrt{2\varepsilon} i}_C y$

$$\rightarrow A + cy + O(y^2) \quad (\text{p. 192 above 7.3.10})$$

Combine I and II for general  $y$

$$\psi(y) = \frac{u(y)}{e^{y^2/2}}$$

$u(y) \rightarrow A + cy \quad y \rightarrow 0$   
 $\rightarrow y^m \quad |y| \rightarrow \infty$

only - solution; other solution blows up as  $|y| \rightarrow \infty$ ; not a bound state.

$$\frac{d\psi}{dy} = u' e^{-y^2/2} - y u e^{-y^2/2}$$

$$\frac{d^2\psi}{dy^2} = u'' e^{-y^2/2} - 2y u' e^{-y^2/2} + y^2 u e^{-y^2/2} - u e^{-y^2/2}$$

$$\frac{d^2\psi}{dy^2} + (2\varepsilon - y^2)\psi = (u'' - 2yu' + y^2u + 2\varepsilon u - y^2u - u)e^{-y^2/2}$$

$$= (u'' - 2yu' + (2\varepsilon - 1)u)e^{-y^2/2} = 0$$

or  $u'' - 2yu' + (2\varepsilon - 1)u = 0$  7.3.11 p.192

another ansatz consistent with the limiting behavior of  $u(y)$ :

$$u(y) = \sum_{n=0}^{\infty} C_n y^n$$

$y \rightarrow 0, u \propto A + cy$  gives  $n=0$  lower lim.

$$u'(y) = \sum_{n=0}^{\infty} C_n n y^{n-1} = \sum_{n=1}^{\infty} C_n n y^{n-1}$$

$y \rightarrow \infty, n \rightarrow \infty$  OK, although  $n < \infty$  OK too.

$$yu'(y) = \sum_{n=0 \text{ or } 1}^{\infty} C_n n y^n$$

$$u''(y) = \sum_{n=0}^{\infty} C_n n(n-1) y^{n-2} = \sum_{n=0 \text{ or } 1 \text{ or } 2}^{\infty} C_n n(n-1) y^{n-2}$$

"Index Shift": since  $y^0$  (when  $n=2$ ) is first power, set

$$m = n - 2$$

$$m + 1 = n - 1$$

$$m + 2 = n$$

$$u''(y) = \sum_{m=0}^{\infty} C_{m+2} (m+2)(m+1) y^m$$

but now  $m$  is just a symbol; let's recycle  $n = m$  (not the old  $n$ !)

$$u''(y) = \sum_{n=0}^{\infty} C_{n+2} (n+2)(n+1) y^n$$

plug into 7.3.11:

$$\underbrace{\sum_{n=0}^{\infty} C_{n+2}(n+2)(n+1)y^n}_{u''(y)} - \underbrace{2\sum_{n=0 \text{ or } 1}^{\infty} C_n n y^n}_{-2y u'(y)} + (2\varepsilon - 1) \underbrace{\sum_{n=0}^{\infty} C_n y^n}_{u(y)} = 0$$

$$\sum_{n=0}^{\infty} \left\{ C_{n+2}(n+2)(n+1) - 2C_n n + (2\varepsilon - 1)C_n \right\} y^n = 0$$

this term must zero for all n.

so, 

$$C_{n+2} = \frac{(2n+1-2\varepsilon)}{(n+2)(n+1)} C_n$$
7.3.14  
p. 193

means that given  $C_0$ , can compute  $C_2, C_4, C_6, \dots$

given  $C_1$ , can compute  $C_3, C_5, C_7, \dots$

So, given 3 numbers:  $\varepsilon, C_0, C_1$ , can compute  $u(y)$ :

$$u(y) = C_0 \left[ 1 + \frac{(1-2\varepsilon)}{2} y^2 + \frac{(5-2\varepsilon)(1-2\varepsilon)}{4 \cdot 3 \cdot 2} y^4 + \frac{(9-2\varepsilon)(5-2\varepsilon)(1-2\varepsilon)}{6 \cdot 5 \cdot 4 \cdot 3 \cdot 2} y^6 + \dots \right]$$

$$+ C_1 \left[ y + \frac{(3-2\varepsilon)}{3 \cdot 2} y^3 + \frac{(7-2\varepsilon)(3-2\varepsilon)}{5 \cdot 4 \cdot 3 \cdot 2} y^5 + \frac{(11-2\varepsilon)(7-2\varepsilon)(3-2\varepsilon)}{7 \cdot 6 \cdot 5 \cdot 4 \cdot 3 \cdot 2} y^7 + \dots \right]$$

notice now... the factor  $\frac{y^n}{n!}$

this means that, if the series really is infinite (doesn't terminate), then

$$u(y) \xrightarrow{y \rightarrow \infty} e^{y^2}$$

$$\text{but then } \psi(y) = u(y)e^{-y^2/2} \rightarrow e^{+y^2/2}$$

and, that is not a bound state!

For bound states, then, the series must terminate, which means that for some integer  $n$ , must have:

$$2n+1 - 2\varepsilon_n = 0$$

$$\boxed{\varepsilon_n = n + \frac{1}{2}}$$

p. 194  
7.3.18

$$\boxed{E_n = \varepsilon_n \cdot \hbar\omega = (n + \frac{1}{2})\hbar\omega}$$

7.3.20

Eigenfunctions (omitting  $e^{-y^2/2}$ )  $\downarrow$  convention.

$$C_0 \neq 0 \quad n=0 \quad u_0(y) = C_0 = 1 \quad E_0 = \frac{1}{2}\hbar\omega$$

$$C_1 = 0 \quad n=2 \quad u_2(y) = C_0[1 - 2y^2] \quad E_2 = \frac{5}{2}\hbar\omega$$

$$\text{even} \quad C_0 = -2 \cdot 1 \text{ (convention)}$$

$$n=4 \quad u_4(y) = C_0[1 - 4y^2 + \frac{4}{3}y^4] \quad E_4 = \frac{9}{2}\hbar\omega$$

$$C_0 = -4 \cdot 3 \text{ (convention)}$$

with these conventions, called "Hermite Polynomials"

$$C_0 = 0 \quad n=1 \quad E_1 = \frac{3}{2} \hbar \omega \quad U_1(y) = c_1 y$$

$$C_1 \neq 0 \quad n=3 \quad E_3 = \frac{7}{2} \hbar \omega \quad U_3 = c_1 (y - \frac{2}{3} y^3)$$

coefficient of highest power  
 $n$  is  $2^n$

$$c_1 = +2 \text{ (convention)}$$

$$c_1 = -4.3 \text{ (convention)}$$

Putting things together

$$\Psi_E(x) = \Psi_{(n+\frac{1}{2})\hbar\omega}(x) = \Psi_n(x)$$

$$x = by \quad y = \frac{1}{b} x$$

with normalizations chosen above,

$$\Psi_n(x) = \left( \frac{m\omega}{\pi \hbar 2^{2n} (n!)^2} \right)^{1/4} e^{-\left(\frac{m\omega^2 x^2}{2\hbar}\right)} H_n \left( \left(\frac{m\omega}{\hbar}\right)^{1/2} x \right)$$

normalization,  
not computed here  
(will get later)

$e^{-y^2/2}$

earlier called  $U_n(y)$

$$\int_{-\infty}^{\infty} H_n(y) H_{n'}(y) e^{-y^2} dy = \sqrt{\pi} 2^n n! \delta_{nn'}$$

"orthogonal polynomials" under weight  $e^{-y^2}$

QSHO comments (p.197 + on)

- ① "macroscopic" oscillator: say  
 $m = 2 \text{ grams} = 2 \cdot 10^{-3} \text{ kg}$   
 $\omega = 1 \text{ rad/sec}$   
 $X_0 = \text{max displacement} = 1 \text{ cm} = 10^{-2} \text{ m}$

②  $E = \frac{1}{2} k X_0^2 = \frac{1}{2} m \omega_0^2 X_0^2$   
 $= \frac{1}{2} \cdot 2 \cdot 10^{-3} \cdot 1^2 \cdot 10^{-4} \text{ Joules}$   
 $\approx 10^{-7} \text{ Joules.}$

③  $\hbar \omega = (6.6 \cdot 10^{-34} \text{ J}\cdot\text{s}) \cdot 1 \frac{1}{\text{s}}$   
 $\approx 6.6 \cdot 10^{-34} \text{ Joules.}$

④  $n \approx \frac{E}{\hbar \omega} - \frac{1}{2} \approx \frac{10^{-7}}{6.6 \cdot 10^{-34}} - \frac{1}{2} \approx 10^{26}$   
 $n \approx 10^{26} \gg 1$  not quantum at all

⑤ This problem's results are adapted to describe lots of boson's quantum properties: photons, phonons, ...

⑥ "Zero point energy" of  $\frac{1}{2} \hbar \omega$ .  
 Not 0 by uncertainty principle (p.199)



④  $n$  even  $\psi_n(-x) = \psi_n(x)$  (even powers of  $x, y$ )  
 $n$  odd  $\psi_n(-x) = -\psi_n(x)$  (odd powers of  $x, y$ )

define "Parity Operator"  $\tilde{\Pi}$

By definition  $|\psi\rangle \equiv \psi(x)$

$\tilde{\Pi}|\psi\rangle \equiv \psi(-x)$

a.k.a.  $\tilde{\Pi}\psi(x) = \psi(-x)$

The eigenstates of the QSHO are also eigenstates of  $\tilde{\Pi}$  with eigenvalues  $\pm 1$ :

$\tilde{\Pi}\psi_n(x) = (-1)^n \psi_n(x)$

⑤ QSHO states tunnel past the classical turning points

⑥  $P_n(x) = |\psi_n(x)|^2$  unlike classical, where

$P(x) \propto dt \propto \frac{dx}{v(x)} \propto \frac{1}{w(x_0^2 - x^2)^{1/2}}$

but as  $n \rightarrow \infty$  the average value of  $P_n(x)$  does resemble classical probability.